

## Pressure profile shape constancy in L-mode stellarator plasmas

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It is well known that the temperature profiles in a tokamak are self-consistent [1-3]. For the stellarators there is no temperature profile consistency [4]. The pressure profile consistency concept works in tokamaks [5]. There was found one example of the pressure profile stiffness in the pellet injection versus gas puff experiments in LHD [6]. The paper is dedicated to the investigation of this concept for the stellarators. Plasma temperature and density profile evolution during NBI heating of on- and off- axis ECRH heated plasma on TJ-II [7], for ECRH power scan on W7-AS [8] and CHS [9], high  $T_i$  mode on CHS [10], on- and off- axis ECRH on W7-AS [4] and gas puffing on ATF [11] were observed (see Table 1).

In the TJ-II stellarator, NBI heating ( $P_{\text{NBI}} = 300$  kW) of the target ECRH plasma ( $P_{\text{ECRH}} = 300$  kW) leads to dramatic changes of the plasma density and temperature.  $n_e$  and  $T_e$  profile evolution measured by high resolution Thomson Scattering diagnostics is shown in Fig.1 [7]. The values varied up to an order of magnitude, ( $0.3 < n_e(0) < 6 \times 10^{19} \text{ m}^{-3}$ ,  $0.2 < T_e(0) < 1 \text{ keV}$ ), the profiles varied from hollow to peaked (density), and from peaked to flat (electron temperature). In spite of the large difference in  $n_e$  and  $T_e$  profiles in the analyzed regime, their product, the plasma pressure  $P_e$ , presents much stronger profile resilience in the confinement zone of the plasma column. It was found that the normalized pressure profiles  $P^{\text{norm}} = P(\rho)/P(0)$  are much less scattered in comparison with plasma  $n_e$  and  $T_e$  profiles, see Fig. 2.

In the CHS experiments with on-axis  $P_{\text{EC}} = 150 \div 215$  kW and density  $n_e = (0.47 \div 0.95) \times 10^{19} \text{ m}^{-3}$  variation [9], the similar behavior was found: the increase of  $T_e$  was accompanied by the decrease of  $n_e$ , leaving  $P^{\text{norm}}(\rho)$  practically unchanged. In the experiments with the standard major plasma axis ( $R_{\text{ax}}=92.1\text{cm}$ ,  $n_e(0)= 4 \times 10^{19} \text{ m}^{-3}$ ,  $T_e(0) \sim T_i(0)=250 \text{ eV}$ ), and optimized one ( $R_{\text{ax}}=87.7\text{cm}$ ,  $n_e(0)= 2 \times 10^{19} \text{ m}^{-3}$ ,  $T_e(0)=200 \text{ eV}$ ,  $T_i(0)=130 \text{ eV}$ ) [10], the same tendency was found,  $P^{\text{norm}}$  remains almost constant. In high  $T_i$  mode ( $n_e(0)= 1.4 \times 10^{19} \text{ m}^{-3}$ ,  $T_e(0)=700 \text{ eV}$ ,  $T_i(0)=1 \text{ keV}$ ) [10],  $P^{\text{norm}}$  almost coincides with the rest of discussed CHS profiles.

In W7-AS experiment with on-axis  $P_{\text{EC}}$  variation from 0.2 to 0.8 MW at almost the same density  $n_e \sim 2 \times 10^{19} \text{ m}^{-3}$  [8], the similar behavior was found: increase of  $T_e$  was accompanied

by a decrease of  $n_e$ , remaining  $P^{norm}(\rho)$  practically unchanged. In another experiment, on- and off- axis ECRH alternate with corresponding  $T_e$  and  $n_e$  variations [4]. It is highlighted in [4] that during off-axis ECRH ( $\rho=0.6$ ) the central density is peaking without an additional particle source. It led to almost unchanged  $P^{norm}(\rho)$ .

At the gas puffing in the ATF [11], the plasma density rise was accompanied by the concordant decay of  $T_e$ , again remaining  $P^{norm}(\rho)$  practically unchanged.

Despite the difference in the magnetic configurations (heliac TJ-II, heliotron CHS, torsatron ATF, optimized W7-AS), a remarkable similarity is seen in the normalized pressure profiles, in other words, the normalized pressure profile has universal shape for normal confinement (L-mode) in all the observed experiments (see Fig. 2). The universal profiles can be fitted by a quasi-linear function in the confinement zone ( $0.2 < \rho < 0.8$ ),  $P_0^{-1}dP/d\rho \sim \text{const} = k =$

$P_0^{-1}\Delta P_{linear}/\Delta\rho_{linear}$ ,  $\Delta\rho_{linear}$  is the radial extension, where profile has almost linear shape. In the observed cases  $\Delta\rho_{linear} \sim 0.6$ ,  $k \approx 1.3 \pm 0.1$  (in Table 1,  $\Delta k$  is the linear regression error for  $k$ ).

Table 1.

Device	Type	$R$ , m	$a$ , m	$B$ , T	$\nu/2\pi(a)$	$-k$	$\Delta k$	Regime, [Ref.]
TJ-II	heliac	1.5	0.22	1.0	1.6, low shear	1.34 1.39	0.04 0.04	EC-on +NBI [7] EC-off +NBI [7]
W7-AS	modular coils	2.0	0.2 0.16 0.13	2.5	low shear, 0.56	1.19 1.28 1.38	0.02 0.02 0.02	$P_{EC}$ scan, $n_e$ -const [8] EC on-off [4] HDH vs NC [12]
CHS	heliotron	0.92	0.19	0.9 1.9 1.9 1.9	strong shear	1.32 1.39 1.36 1.29	0.02 0.02 0.02 0.02	EC-on, $n_e$ -var [9] EC-on, ITB [10] High $T_i$ , NBI 1MW $R_{ax}$ -var., $n_e$ -var. [10]
ATF	torsatron	2.1	0.27	1.9	shear	1.11	0.02	Gas puff [11]

Contrary to the L-mode, pressure profiles show different shapes during improved confinement modes. An example of the edge transport barrier is shown in Fig. 2b.  $P^{norm}(\rho)$  in HDH confinement mode in W7-AS [12] differs strongly from universal profile, while the reference Normal Confinement profile belongs to the universal one. Strong similarity of the profiles allows us to get any spatial point in the confinement area for the normalization, not only  $\rho=0$ . In case of ITB formation the pressure profiles have clearly two components. The outer component (outside the ITB), normalized at the ITB foot, shows the pressure profiles similarity (the universal profile takes place), while inside the ITB area  $P_{ITBfoot}^{-1}dP/d\rho$  is significantly higher. Fig. 3 shows W7-AS data with on-axis  $P_{EC} = 1.2$  MW, where the

temperature and pressure profiles show the ITB formation at  $\rho \sim 0.25$  [8]. In CHS ITB was obtained for on-axis ECRH ( $B=1.9T$ ,  $P_{EC} = 0.3$  MW at 106GHz) at  $\rho \sim 0.4$  [10]. In both cases

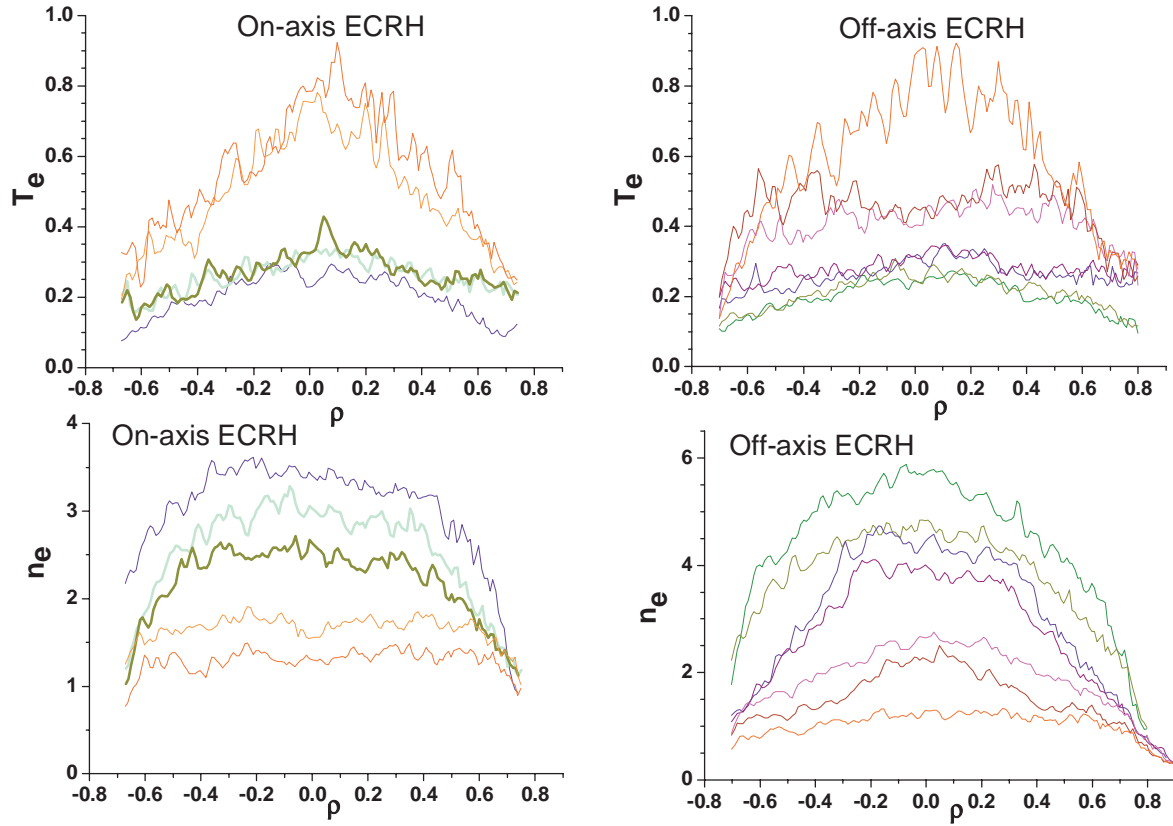


Fig 1.  $T_e$  (upper) and  $n_e$  (lower) profile evolution in the NBI experiments for on axis (left) and off-axis (right) ECRH in TJ-II. Time passes from red to blue curves.

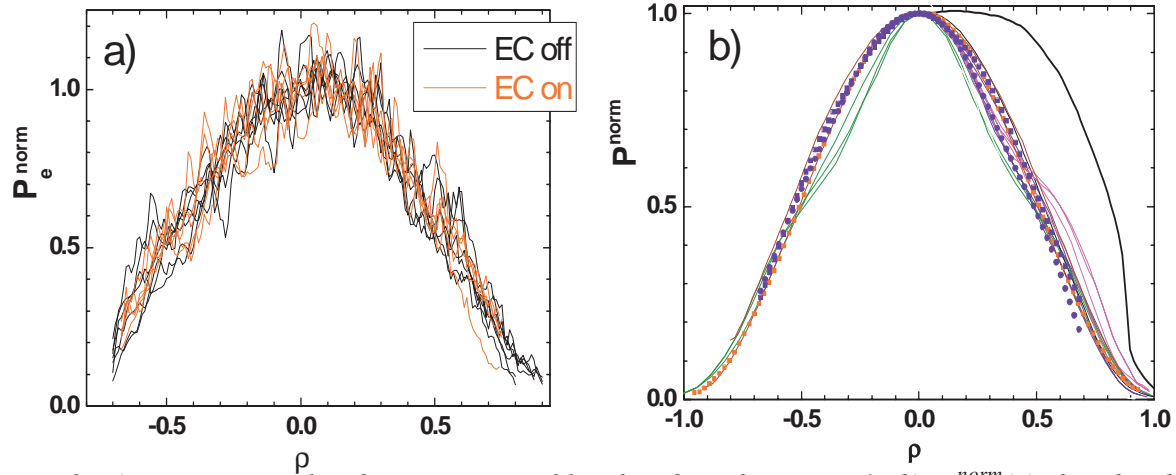


Fig. 2. a) TJ-II normalized pressure profiles for data from Fig.1; b)  $P^{norm}(\rho)$  for the data presented in table forms the universal profile; CHS: red dots – high  $T_i$  mode, brown –  $R_{ax}$  variation, red –  $P_{EC}$  variation; ATF: purple - gas puff; W7-AS: green -  $P_{EC}$ -on scan, blue - EC-on- and off-axis, black – HDH-mode, dark blue – reference Normal Confinement mode; TJ-II: blue dots – EC on-, off-axis. Profiles from all machines and experiments were radially normalized to the actual plasma size,  $P^{norm}(1)=0$ .

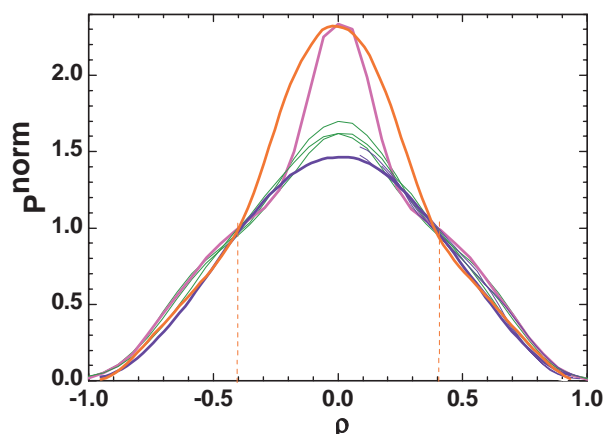


Fig. 3. The universal profile features outside the ITB area. CHS: blue – high  $T_i$  mode, red – ITB; W7-AS: green –  $P_{EC} = 0.2-0.8$  MW, pink – ITB ( $P_{EC} = 1.2$  MW);  $P^{norm} = P/P(\rho_{ITBfoot} = 0.4)$

remarkable similarities between each other in low (TJ-II, W7-AS) and high (CHS, ATF) magnetic shear configurations. So, the universal profile, characterized by  $P_0^{-1}dP/d\rho \sim \text{const} = k$ , was found for L-mode plasmas and outside the ITB area. The other types of profile like in case of LHD [6] may take place for specific plasma conditions. (III) The observation of the universal constant  $k$  in L-mode (e.g. in the absence of strong  $E \times B$  effects) may suggest that turbulence and the associated transport reach some kind of saturation level, that does not depend on the absolute values of  $T_e$  and  $n_e$ , but alternative explanations (e.g. based on the role of atomic physics and links between magnetic configuration and gradient) can not be excluded. The validation of the transport-based hypothesis would require to characterize the link between local gradients and turbulent transport.

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outside the ITB area, the profiles coincide with the universal one.

**Summary.** (I) In spite of large variety in the profiles of plasma electron temperature and density, their product, electron pressure presents the feature of profile constancy in stellarator devices in observed experiments with wide range of the plasma and heating parameters.

(II) In the L-mode plasmas of the medium size stellarators, the pressure profiles show